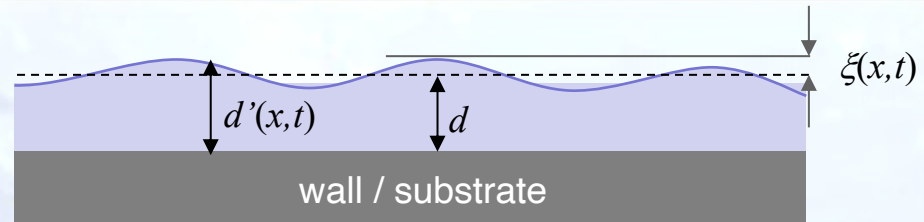




(iii) Third sound

sound propagating in **thin** films



$$d'(x,t) = d + \xi(x,t)$$

mean film thickness

assumptions: thin films $v_n = 0$

$\lambda \gg d \longrightarrow$ motion parallel to substrate in x-direction ($v_y = v_z = 0$)

$\text{grad } T \approx 0$ (questionable ?)

$$\longrightarrow \frac{\partial^2 \xi}{\partial t^2} = f d \frac{\rho_s}{\rho} \frac{\partial^2 \xi}{\partial x^2} \longrightarrow$$

v. d. Waals force

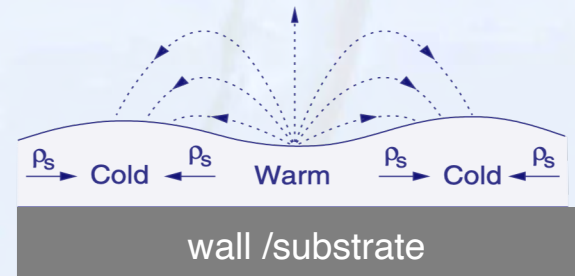
$$v_3^2 = \frac{\rho_s}{\rho} 3gz$$

height over liquid level

problem: evaporation and condensation

\longrightarrow increases the amplitude and changes velocity

can be taken into account



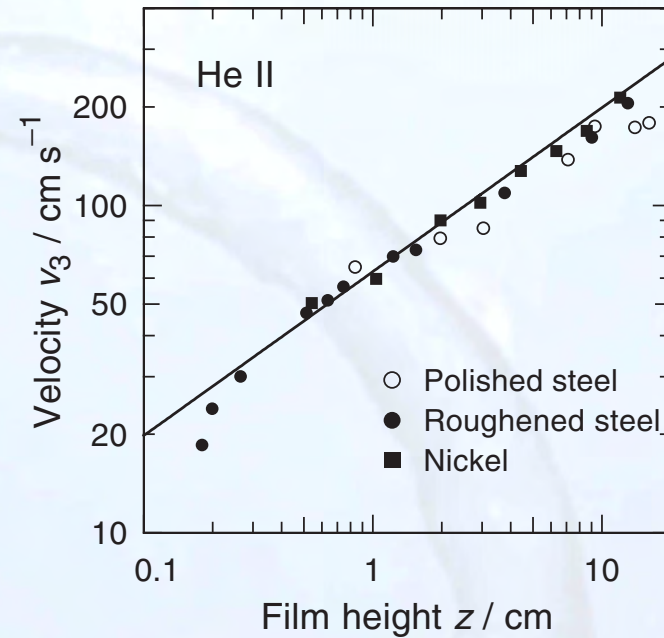
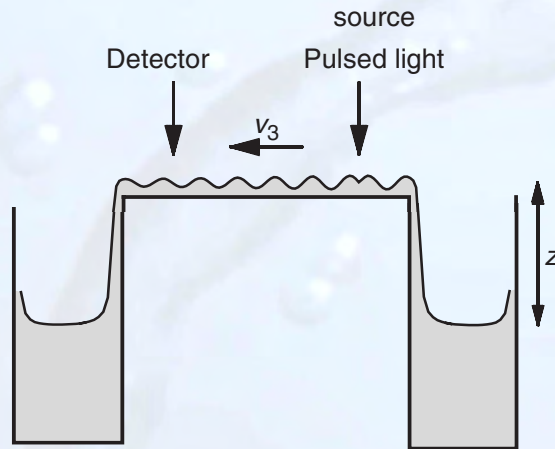
$$v_3^2 \approx \frac{\rho_s}{\rho} 3gz \left(1 + \frac{TS}{L} \right)$$

$$\left. \begin{aligned} TS/L &= 0.01 \text{ at } 1 \text{ K} \\ TS/L &= 0.15 \text{ at } T_\lambda \end{aligned} \right\}$$

not really a problem



3rd sound experiment



Procedure

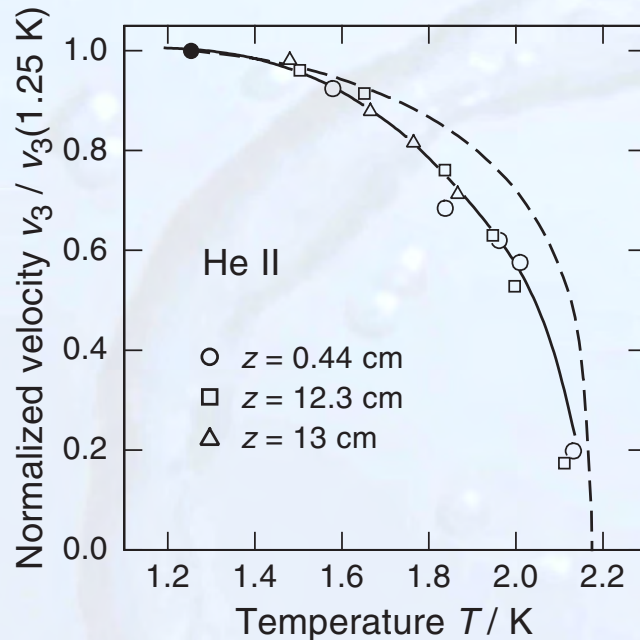
- ▶ periodic local heating
- ▶ Q_s flows to **warm location** → thickness changes
- ▶ surface wave \triangleq **3rd sound**
- ▶ optical detection of thickness

Measurement and results

- ▶ 3rd sound velocity vs. z (log/log plot)
- ▶ different surfaces: v_3 almost independent
- ▶ line \triangleq theory $v_3 \propto \sqrt{z}$
- ▶ good agreement except for very thick films



3rd sound experiment: temperature dependence



Measurement and results

- ▶ 3rd sound velocity vs T
- ▶ points at $T = 1.25$ K normalized to (●)
- ▶ v_3 is rising with decreasing T
- ▶ $T \rightarrow 0$: $v_3 = 1.5$ m/s (very slow)
- ▶ dashed line \triangleq theory $v_3 \propto \sqrt{\rho_s}$
- ▶ systematic deviations: origin unknown, but likely due to generation process

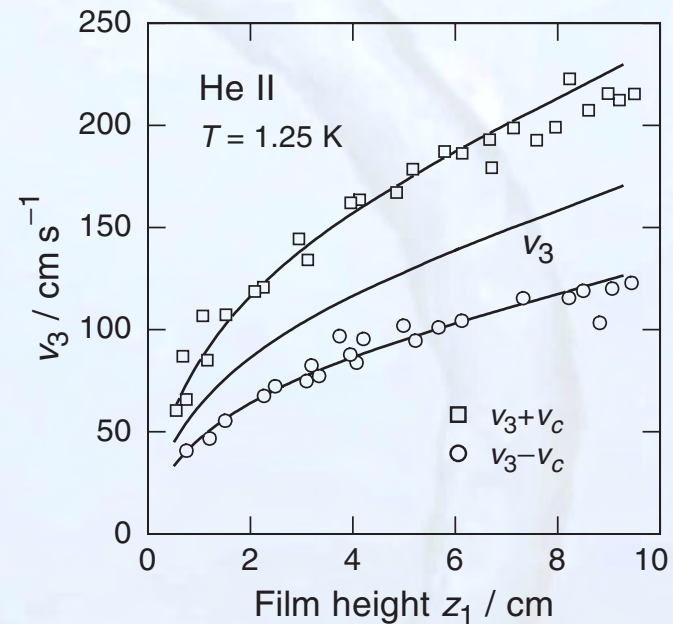
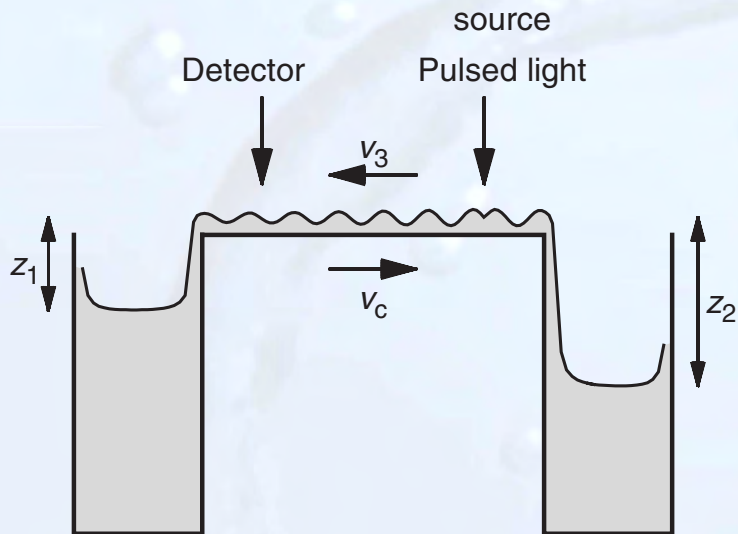
3rd sound in very thin films:

3rd sound propagation can be observed down to 2.1 monolayers

➡ onset of superfluidity



3rd sound in moving films:



3rd sound propagation in moving films \rightarrow Doppler effect

$v_3 \pm v_f$ $\xrightarrow{v_c}$ critical velocity

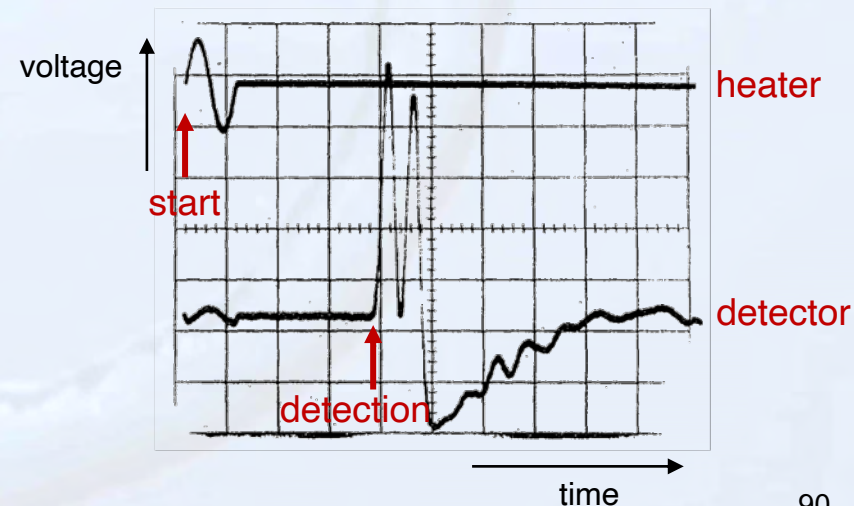
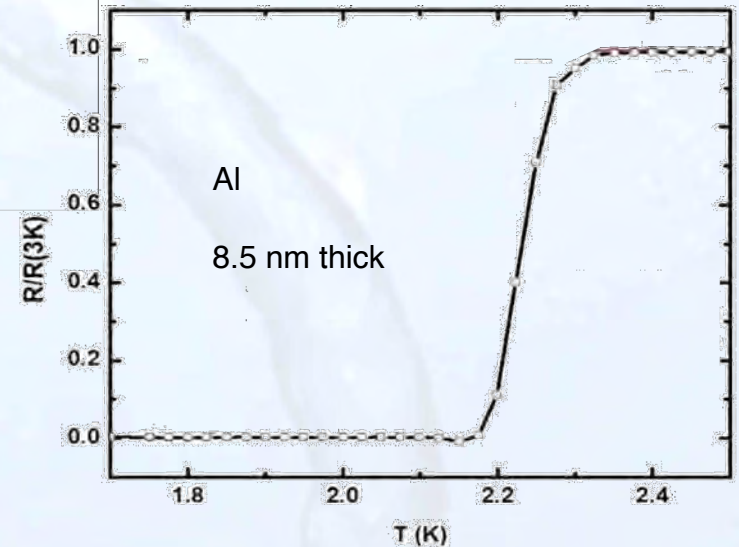
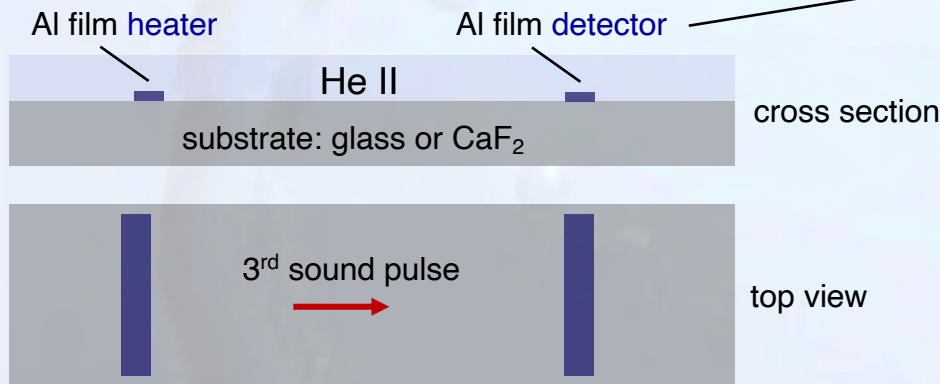


Detection of 3rd sound experiment in ultralow films:

Third Sound and the Healing Length of He II in Films as Thin as 2.1 Atomic Layers*

J. H. Scholtz, E. O. McLean,† and I. Rudnick
University of California, Los Angeles, California 90024
(Received 23 August 1973)

Measurements of the velocity of third sound on films as thin as 2.1 atomic layers yield the healing length of superfluid He II down to temperatures of 0.1 K. It is argued that these films are two-dimensional superfluids. **PRL 32 147 (1974)**



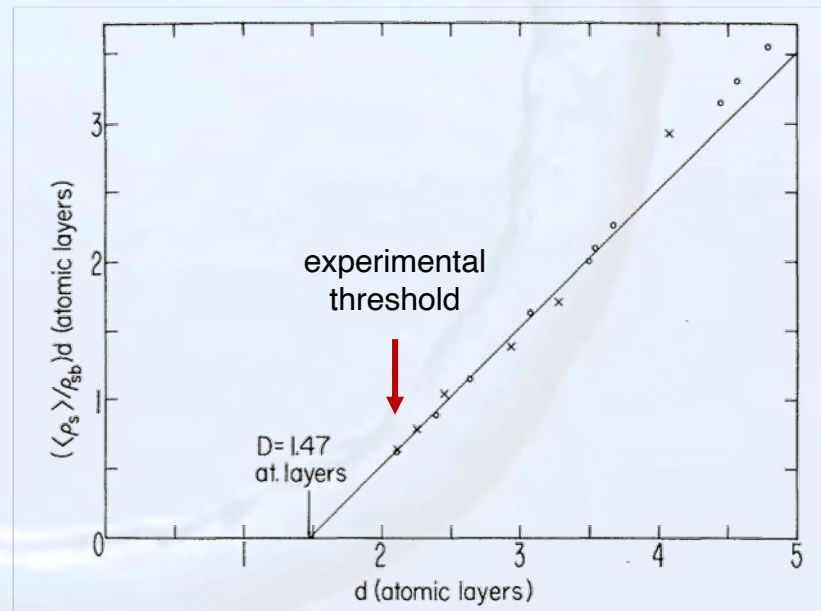
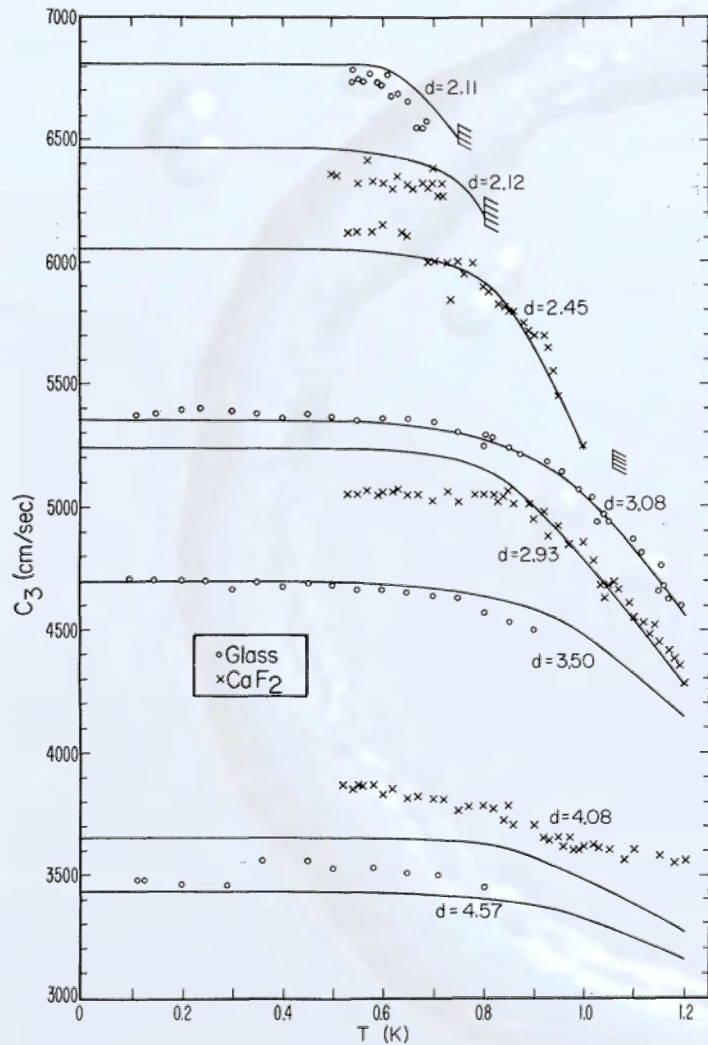
- ▶ time of flight detection of 3rd sound pulse
- ▶ thin film Al heater and detector
- ▶ detector operated at the transition temperature
- ▶ detection of heat pulse associated with 3rd sound pulse
- ▶ very sensitive because of steep transition curve



Experimental results:

for ultrathin films:
$$v_3^2 = \frac{\bar{\rho}_s}{\rho_{s,bulk}} \frac{3RT}{m} \ln \frac{p_0}{p}$$

- ▶ experimental threshold of **2.1 monolayers** independent of substrate
- ▶ **film thickness** determine by **amount of helium** and **surface area**
- ▶ **extrapolation** suggests that **1.57 monolayers** might be the onset threshold





(iv) Fourth sound

sound propagation in fine powders / slits $v_n \approx 0$

→ oscillations in total density, in ratio of superfluid to normalfluid density,
in pressure, in temperature, in entropy

$$v_4^2 = \frac{\rho_s}{\rho} v_1^2 \left[1 + \underbrace{\frac{2ST}{\rho C_p} \left(\frac{\partial \rho}{\partial T} \right)_p}_{\ll 1} \right] + \frac{\rho_n}{\rho} v_2^2$$

$$v_4 \approx \sqrt{\frac{\rho_s}{\rho} v_1^2 + \frac{\rho_n}{\rho} v_2^2} \approx \sqrt{\frac{\rho_s}{\rho} v_1^2}$$

5th sound4th sound generation like for 1st sound, but $v_n \approx 0$



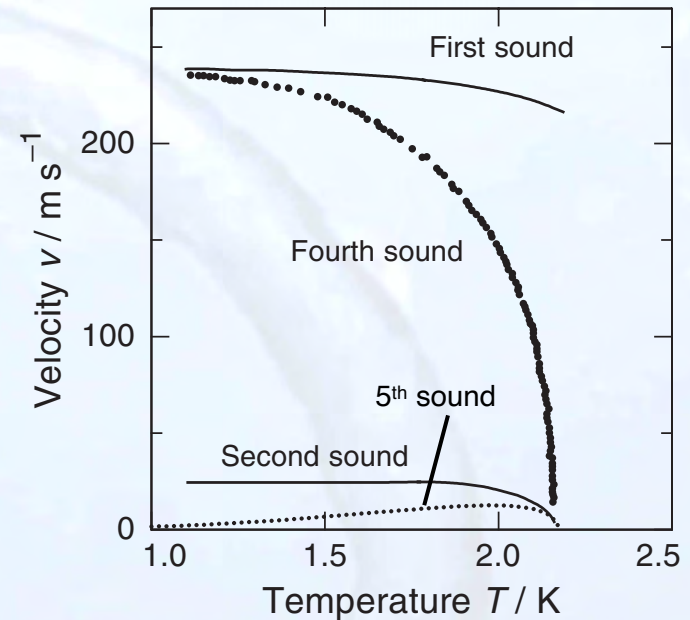
4th sound experiments

4th sound generation like for 1st sound, but $v_n \approx 0$

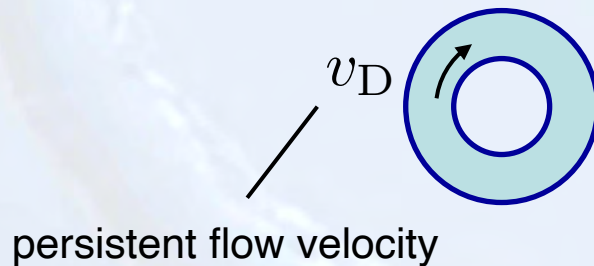
$T \rightarrow 0$ $v_4 = v_1 \approx 238 \text{ m/s}$, since $\varrho_s = \varrho$

$T = T_\lambda$ $v_4 = 0$

$$v_4 \approx \sqrt{\frac{\varrho_s}{\varrho} v_1^2 + \frac{\varrho_n}{\varrho} v_2^2}$$



Persistent flow and 4th sound



$$v_4 \approx v_{4,0} \pm \frac{\varrho_s}{\varrho} v_D$$

coupling of a compression wave
to second sound



Einstein 1924
Bose 1925
London 1938

Basic idea of Fritz London:

dissipation-less motion \longleftrightarrow macroscopic wave function

a) Ideal Bose gas

non-interacting Bose gas (rough approximation for liquid He)

let's consider: 1 cm³ cube of liquid ⁴He $\triangleq 10^{22}$ atoms with mass m

eigenstates for **free particles** in a cube:

$$E_n = \frac{\hbar^2}{2m} \left(\frac{\pi}{L} \right)^2 n^2 \quad \text{with} \quad n^2 = n_x^2 + n_y^2 + n_z^2$$

$T = 0 \longrightarrow$ **all** atoms are in the ground state E_{111} **trivial !**

But at finite temperatures?



consider **energy difference** between **ground state** and **first excited state**

$$\Delta E/k_B = (E_{211} - E_{111})/k_B \approx 2 \times 10^{-14} \text{ K}$$

↪ if **Boltzmann** statistics would hold ➡ **no condensate at 1 K!!!**

however, **Bose-Einstein distribution** is relevant here

$$f(E, T) = \frac{1}{e^{(E-\mu)/k_B T} - 1} ,$$

/

chemical potential $\mu = \frac{\partial F}{\partial N}$

what we know: $\mu < E_{111}$ ➡ otherwise, **negative** occupation
 $\mu \neq 0$ ➡ since particle number **conserved**



Occupation of ground state $E_{111} = 0$

$$f(0, T) = \frac{1}{e^{-\mu/k_B T} - 1} \quad \longrightarrow \quad \text{occupation depends critically on } \mu$$

$$f(0, T \rightarrow 0) \rightarrow \infty \quad \text{if } \mu \rightarrow 0 \text{ faster than } T \rightarrow 0 \quad \left(\frac{1}{e^0 - 1} \rightarrow \infty \right)$$

What is the temperature dependence of $\mu(T)$?

for this let us consider a **real**, but **non-interacting** gas

$$\mu = -k_B T \ln \left(\frac{V_A}{V_Q} \right)$$

quantum volume $V_Q = \left(\frac{h}{\sqrt{2\pi m k_B T}} \right)^3 = \lambda_B^3$

thermal de Broglie wavelength

For $^4\text{He} \longrightarrow \lambda_B^3 = (8.7 \text{ \AA})^3 \text{ at } 1 \text{ K}$

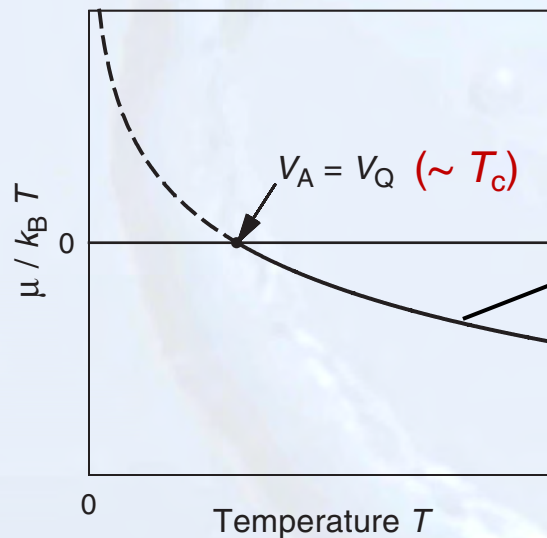
$V_A = V/N = (3.8 \text{ \AA})^3 \text{ in comparison}$



for $T \rightarrow 0 \rightarrow V_Q \rightarrow \infty$

↪ at sufficiently low temperatures $V_A = V_Q$!

$$\ln \left(\frac{V_A}{V_Q} \right) \rightarrow 0 \rightarrow \mu \rightarrow 0$$

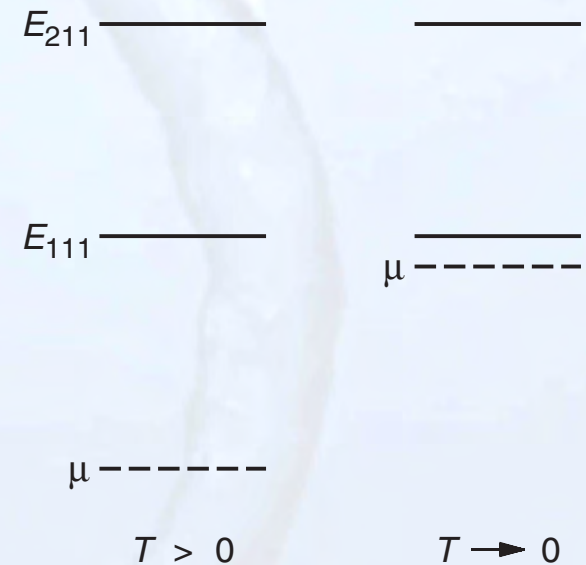


classical regime

μ is negative

He gas at 300 K:

$$\mu/k_B = -3800 \text{ K}$$



↪ this means $|\mu|$ becomes **smaller** than $\Delta E/k_B = (E_{211} - E_{111})$ at finite T !



Calculation of μ : how large is μ at 1K? (revers argument)

for $T \rightarrow 0 \rightarrow f_{111} \rightarrow N$

$$\lim_{T \rightarrow 0} f(0, T) = N_0(T) = \lim_{T \rightarrow 0} \left(\frac{1}{e^{-\mu/k_B T} - 1} \right)$$

$E_{111} = 0$, ground state

$$\approx \lim_{T \rightarrow 0} \left(\frac{1}{1 - \mu/(k_B T) + \dots - 1} \right) \approx -\frac{k_B T}{\mu}$$



$$\mu = -\frac{k_B T}{N_0}$$

close to $T = 0$



at $T = 1 \text{ K} \rightarrow \mu/k_B \approx 10^{-22} \text{ K}$





Calculation of N_0 and N_e :

N_e :
number of particles in excited states

$$\begin{aligned} \sum_i f(E_i, T) &= N = N_0(T) + N_e(T) \\ &= N_0(T) + \int_0^\infty D(E) f(E, T) dE \end{aligned}$$

$D(E)$: density of states for **free** particles without $D(0)$

density of states for **free** particles $E_k \propto k^2$

$$D(E) = \frac{V(2m)^{3/2} \sqrt{E}}{4\pi^2 \hbar^3}$$

with $E/k_B T = x$ **and** $|\mu| \ll \Delta E \longrightarrow \exp[(E - \mu)/k_B T] \approx \exp(E/k_B T)$

$$\begin{aligned} \longrightarrow N &= N_0 + \frac{V}{4\pi^2} \left(\frac{2m}{\hbar^2} \right)^{3/2} (k_B T)^{3/2} \underbrace{\int_0^\infty \frac{\sqrt{x}}{e^x - 1} dx}_{\Gamma(5/2) \times \zeta(5/2) \approx 1.783} \end{aligned}$$



with $V_Q = \left(\frac{h}{\sqrt{2\pi m k_B T}} \right)^3 = \lambda_B^3$

$\curvearrowright N \approx N_0 + 2.6 \frac{V}{V_Q}$

$$N_0 = N - 2.6 \frac{V}{V_Q}$$

Interpretation

as long as $2.6 \frac{V}{V_Q} \ll 10^{22}$, which means that the de Broglie **wavelength** is **significantly larger** as an **atom** \rightarrow **condensation** !

factor $\sqrt[3]{2.6} = 1.37$

- ▶ $T = 0 \rightarrow N_0 = N$ trivial !
- ▶ $0 < T < T_c \rightarrow N_0$ still **macroscopically large**!
- ▶ $N_e \triangleq$ **normalfluid** component

comment:

λ_B^3 must not be as large as the vessel as proposed by London